## PHYSICS OF ELEMENTARY PARTICLES AND ATOMIC NUCLEI. EXPERIMENT

# On the Study of Antiprotons Yield in Hadronic Collisions at NICA SPD

V. Alexakhin<sup>a, \*</sup>, A. Guskov<sup>a, \*\*</sup>, Z. Hayman<sup>b, \*\*\*</sup>, R. El-Kholy<sup>b, \*\*\*\*</sup>, and A. Tkachenko<sup>a, \*\*\*\*\*</sup>

<sup>a</sup> Joint Institute for Nuclear Research (JINR), Dubna, 141980 Russia
<sup>b</sup> Astronomy Department, Faculty of Science, Cairo University, Giza, 12613 Egypt
\*e-mail: vadim.alexakhin@cern.ch
\*\*e-mail: alexey.guskov@cern.ch
\*\*\*e-mail: zmhayman@cu.edu.eg

\*\*\*\*e-mail: relkholy@sci.cu.edu.eg

\*\*\*\*\*e-mail: avt@jinr.ru

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**Abstract**—The study of antiproton yield in p-p and p-d collisions is important for the astrophysical search for dark matter consisting of Weakly Interacting Massive Particles. Refinement of the production cross section, angular and momentum spectra of produced antiprotons in a wide collision energy range could help to treat the results obtained by the AMS-02 and PAMELA orbit spectrometers. In this paper we present a detailed Monte Carlo study of possible measurements at the planned SPD experiment at the NICA collider that is under construction at JINR (Dubna).

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## INTRODUCTION

Dark Matter (DM) is thought to compose more than 24% of the matter-energy content of the Universe [1]. While observations have not yet revealed the origin or nature of DM, there are many theories trying to infer both from available information. One of the most common theories is that the galactic dark halos are built up of thermal relics of the Big Bang called Weakly Interacting Massive Particles (WIMPs) [2]. There is a possibility that, even though DM particles do not interact with Standard Model (SM) particles, they do produce SM particles via pair-annihilation or decay. Powerful experiments, such as AMS-02 and PAMELA, have been trying to detect anomalies in the spectrum of rare components of Cosmic Rays (CRs), especially antiparticles like antiprotons, that would indicate an exotic signal that could be coming from DM. However, conclusions are hindered by uncertainties: mainly, those that surround cross-sections of secondary CR production, propagation parameters in the galactic environment, the spectrum of higherenergy primary CRs, and the modulation associated with solar activity [3].

To be able to detect anomalous components in CRs, it is crucial to measure the production of these components from conventional astrophysical sources with the maximal possible precision. For the second-ary antiproton background, the sources are proton-nucleus, nucleus-proton, and nucleus-nucleus colli-

sions of CRs on the Interstellar Medium (ISM), which is the dust and gas in ionic, atomic, and molecular form, existing in the galactic environment among the stellar systems. However, almost all secondary antiprotons are a result of collisions of light nuclei. Despite the AMS-02 measuring the antiproton flux with unprecedented accuracy of a few percents [4], no conclusion can be made yet about exotic components due to several sources of uncertainty. The largest source of uncertainties, which can vary from 20 to 50% depending on the energy, is uncertainties on antiproton-production cross sections [3].

In addition to the lack of data of antiproton production in collisions involving helium, data of antiproton production in proton-proton collisions are scarce. Moreover, in most of the existing cross-section datasets for proton-proton collisions - which can be found in [5]—no feed-down contribution has been applied to account for antiproton production via the decay of intermediate hyperons; namely,  $\overline{\Lambda}$  and  $\overline{\Sigma}^-$ ; a process that gains significance with increasing energy [6]. In order to keep up with the AMS-02 accuracy and energy range, new accurate measurements of antiproton production are essential. This would enable comparison of data with theoretical models; and, eventually, coming to a conclusion about any exotic signals. These measurements need to cover the proton-beam kinetic-energy range from 1 GeV to 6 TeV, and the pseudorapidity range from 2 to 8 [7].

It is currently planned to measure the antiprotonproduction cross-sections of p-p and p-He collisions in fixed-target mode by the COMPASS Collaboration [8]. However, the collider mode and  $4\pi$ -coverage of the Spin Physics Detector (SPD) detector [9] at the Nuclotron-based Ion Collider fAcility (NICA) [10] will enable measurement of antiprotons yield at similar collision energy ( $\sqrt{s}$ ) but with higher transverse momentum ( $p_T$ ). There is also the possibility of performing proton-deuteron collision studies, which contribute to the secondary antiproton flux in CRs as well.

NICA is currently under construction at JINR. It is planned to operate the NICA collider with polarized proton and deuteron beams for spin physics study at the SPD detector. In a previous work [11], we performed a preliminary study of antiproton production at the SPD. Monte Carlo simulations generated with Pythia8 have shown that the momenta of the antiprotons produced in p-p collisions, at a center-of-mass (CM) energy  $\sqrt{s} = 26$  GeV, would be low enough to identify with the time-of-flight method. They have also shown that with a time resolution of about 100 ps and a flight base of 2 m,  $K^-/\overline{p}$  separation can be achieved for antiprotons with momenta up to 4 GeV, which would include most of the antiproton yield.

In this paper, we study with more detail the antiproton production in hadronic collisions with Monte Carlo simulations specific to the SPD setup using the SPDROOT toolkit [12].

#### 1. THE SPIN PHYSICS DETECTOR

The Spin Physics Detector is planned as a multipurpose universal  $4\pi$  detector. It will be able to operate in the polarized p-p collision mode with CM energy up to 27 GeV and luminosity up to about  $10^{32}$  cm<sup>-2</sup> s<sup>-1</sup>. Another mode would be the polarized d-d collisions at energy up to 13.5 GeV and a luminosity one order lower. Asymmetric p-d collisions are also discussed [9]. The detector consists of the barrel part and two end-caps. The tracking system of the SPD includes the silicon-based vertex detector surrounded by the main tracker, using gas-filled drift straw-tubes as the basic detection element with a spatial resolution of about 100 µm. The magnetic field (a few configurations are currently under consideration) provides charged-particle momentum measurement with a typical resolution of about 2% at a transverse momentum of 1 GeV/c. The time-of-flight (TOF) system, with a time resolution below 100 ps, provides identification of secondary hadrons in a wide kinematic range. The shashlyk-type electromagnetic calorimeter is responsible for the photon reconstruction and identification of electrons and positrons. The muon system (RS) performs the advanced muon/hadron separation via comprehensive pattern recognition and matching of the track segments to the tracks in the inner part of the detector [9].

This Monte Carlo study simulation was performed in order to estimate the capability of the Spin Physics Detector for the precision measurement of antiproton yield in hadronic collisions at NICA. The study was done using the SPDROOT toolkit, a ROOT- and GEANT4-based software developed for SPD using the FAIRROOT framework [13]. A simplified description of SPD with a quasi-solenoidal magnetic field along the beam axis (B = 0.8 T) was used. The time resolution of the TOF system used for antiproton identification was supposed to be 70 ps. PYTHIA8 [14] was used as a generator of primary p-p interactions. Since the geometry of the detector is not yet fixed all the results presented in the paper should be treated as indicative.

## 2. DIRECT PRODUCTION OF ANTIPROTONS IN p-p COLLISIONS

The inclusive cross section of antiproton production in p-p collisions in the discussed kinematic range is on the level of a few millibarns [11]. That corresponds to an antiproton production rate on a level  $> 10^5$  s<sup>-1</sup>. So, the accuracy of the proposed measurements is not limited by the statistical uncertainty. The phase space in terms of the variable  $x_R$  and the transverse momentum  $p_T$  of produced antiprotons covered by SPD at  $\sqrt{s} = 13$  and  $\sqrt{s} = 26$  GeV is shown in Fig. 1. Here  $x_R$  is defined as the ratio of antiproton energy to its maximal possible energy for a particular  $\sqrt{s}$ . The geometrical acceptance of the tracking system that affects the track reconstruction efficiency and energy losses of the produced antiprotons, and the magnetic field preventing low-energy particles from reaching the TOF system are taken into account. An antiproton produced at the interaction point was considered detected if it reached the TOF system in the barrel or an end-cap. The radius of the barrel part of the TOF system and the distance from the center of the detector to the end-cap parts are assumed to be 2.0 m each. The detection efficiency as a function of  $\overline{p}$ 

momentum for 
$$\sqrt{s} = 26$$
 GeV is presented in Fig. 2a.

Antiproton identification is the most critical issue for the proposed measurements. The situation is worsening with the presence of only one detector plane for the time-of-flight measurement and with the relatively-low multiplicity of secondary charged particles. The algorithm described in [15] was applied for events with track multiplicities above 3. For each track in the event, three possible hypotheses were assumed:  $\pi$ , *K* and *p*. The  $\chi^2$  minimization of the time mismatch at the interaction point was performed over all possible combinations of individual-track hypotheses. As a



Fig. 1. The kinematic range accessible by p-p collisions at the SPD detector, where the red points represent a CM energy of 13 GeV and the magenta points represent a CM energy of 26 GeV, superimposed on the required measurement range omatch the precision level of the AMS-02 measurements [7].



Fig. 2. (a) Detection efficiency as a function of  $\overline{p}$  momentum for p-p collisions at  $\sqrt{s} = 26$  GeV. (b)  $t_0$  reconstruction accuracy as a function of the number of charged tracks in the event.

result of such minimization, for each event the interaction time  $t_0$  and the most probable particle IDs for each track could be found. Figure 2b shows the resolution of  $t_0$  reconstruction as a function of the number of charged tracks in the event. One can see that the resolution for  $t_0$  becomes better with increasing number of tracks involved into the fitting procedure. Safe  $K^-/\bar{p}$ separation with a purity of antiproton sample on the level of 1% could be provided for particles with

momenta below 3.5 GeV/c.

## 3. ANTIPROTONS FROM ANTI-HYPERONS DECAY

It has been mentioned in [11] that a significant part of antiprotons is produced from the decays of secondary  $\overline{\Lambda}$  and  $\overline{\Sigma}^-$  anti-hyperons. The ratio  $\Delta_{\Lambda}$  of hyperon-induced to directly produced antiprotons can be expressed as

$$\Delta_{\Lambda} = \frac{\overline{\Lambda}}{\overline{p}} \times BF_{\overline{\Lambda} \to \overline{p}\pi^{+}} + \frac{\overline{\Sigma}^{-}}{\overline{p}} \times BF_{\overline{\Sigma}^{-} \to \overline{p}\pi^{0}}, \qquad (1)$$



Fig. 3. Mass reconstruction for  $\pi^-$ ,  $K^-$ , and  $\overline{p}$  using the TOF system. Simulation is performed for  $\sqrt{s} = 26$  GeV.



Fig. 4.  $\overline{\Lambda}/\overline{p}$  ratio in proton-proton collisions as measured by several experiments [6]. The range of  $\sqrt{s}$  accessible at the NICA SPD is shown in red at the lower left part.

where  $\overline{\Lambda}/\overline{p}$  and  $\overline{\Sigma}^{-}/\overline{p}$  are the hyperon to (promptly produced) antiproton ratios,  $BF_{\overline{\Lambda}\to\overline{p}\pi^{+}} = 0.639 \pm 0.005$  and  $BF_{\overline{\Sigma}^{-}\to\overline{p}\pi^{0}} = 0.5157 \pm 0.0003$  are the branching fractions of the corresponding decays. Existing experimental data for the ratio  $\overline{\Lambda}/\overline{p}$  are presented in Fig. 4. The relative uncertainty of this quantity at SPD ener-

gies is about 12%. As for the  $\overline{\Sigma}^{-}/\overline{p}$  ratio, there is not any trustable experimental data on it; and it was estimated roughly based on the symmetry arguments that  $\overline{\Sigma}^{-}/\overline{\Lambda} = 0.33$  [16]. Pythia8 [17] Monte Carlo generator with default settings, for instance, gives that  $\overline{\Sigma}/\overline{\Lambda} \sim 0.30$  at  $\sqrt{s} = 13-26$  GeV [11]. Thus,  $\Delta_{\Lambda} = (0.81 \pm 0.04) \times \overline{\Lambda}/\overline{p}$ .

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Fig. 5. (a) Reconstructed  $\bar{p}\pi^+$  invariant mass spectrum with a  $\bar{\Lambda}$  signal. (b) Reconstructed  $\bar{p}\pi^0$  invariant mass spectrum with a  $\bar{\Sigma}^-$  signal.

Since hyperons have a macroscopic decay length, their contribution can be accessed at the SPD via reconstruction of the two-body decays  $\overline{\Lambda} \rightarrow \overline{p}\pi^+$  and  $\overline{\Sigma}^- \rightarrow \overline{p}\pi^0$  in secondary vertices. The fraction of  $\Lambda$ decaying within the inner part of the SPD setup, averaged over the spectrum, is 93% at  $\sqrt{s} = 26$  while almost all  $\overline{\Sigma}^-$  decay there. The reconstructed invariant mass of  $\overline{p}\pi^+$  from  $\overline{\Lambda}$  decay in secondary V-vertices is shown in Fig. 5a. The Gaussian width of the peak is about 2 GeV/ $c^2$ . It is fully defined by the uncertainty of the momentum measurement in the SPD tracking system.

For determination of the  $\overline{\Sigma}^-$  decay, a secondary vertex should be reconstructed using a track pointing to the primary vertex and associated with  $\overline{\Sigma}^-$  and a track associated with an antiproton. A pair of photons detected in the electromagnetic calorimeter should be assumed to be produced from a  $\pi^0$  decay in the found secondary vertex. The width of the reconstructed  $\overline{\Sigma}^$ peak is mainly defined by the energy resolution of the calorimeter and equals to 19 MeV as it can be seen from Fig. 5b. Both signals,  $\overline{\Lambda}$  and  $\overline{\Sigma}^-$ , are well-distinguished over the background. That inspires a hope to determine the ratios  $\overline{\Lambda}/\overline{p}$  and  $\overline{\Sigma}^-/\overline{p}$  and, finally,  $\Delta_{\Lambda}$ with a minimal uncertainty. Since huge statistics are expected, this uncertainty will be fully defined by possible systematic effects that have to be estimated later.

### 4. SUMMARY

Based on the proposal to use the Spin Physics Detector at the NICA collider as an instrument for the precision measurements of the antiproton yield in

p-p and p-d collisions required by astrophysical searches for dark matter [11], we performed this Monte Carlo study using a preliminary SPD setup geometry. We refined the kinematic coverage of the detector, studied its possibility to identify directly produced antiprotons with the TOF system, and investigated the conditions of reconstructing the  $\overline{\Lambda}$  and  $\overline{\Sigma}^$ hyperon decays. It was shown that the SPD TOF system on the base of 2.0 m can provide purity of  $K^{-}/\overline{p}$ separation better than 99% up to p = 3.5 GeV/c. Accuracy of the interaction time  $t_0$  reconstruction could reach 40–50 ps. The SPD tracking system and electromagnetic calorimeter offer the possibility of effectively reconstructing the decays  $\overline{\Lambda} \rightarrow \overline{p}\pi^+$  and  $\overline{\Sigma}^- \to \overline{p}\pi^0$ , which are the main sources of secondary antiprotons. The setup resolution of the  $\overline{\Lambda}$  and  $\overline{\Sigma}^$ peaks is found to be 2 and 19 MeV, respectively.

We reaffirm our previous conclusions that the supplementary studies at the Spin Physics Detector could make a sizable contribution to the search of physics beyond the Standard Model by precision measurement of antiproton yield in p-p and p-d collisions, direct and the feed-down contributions, in the kinematic range from the threshold to 27 GeV. Taking into account high luminosity and large production cross section, we could expect that the statistical error of such measurement is negligibly small, and the overall precision is fully defined by systematic effects.

#### REFERENCES

 G. Hinshaw et al. (WMAP Collab.), "Nine-year Wilkinson microwave anisotropy probe (WMAP) observations: Cosmological parameter results," Astrophys. J. Suppl. Ser. 208 (2), 19 (2013).

- 2. M. Tanabashi et al. (PDG Collab.), "Review of particle physics," Phys. Rev. D **98** (3) (2018).
- G. Giesen, M. Boudaud, Y. Genolini, V. Poulin, M. Cirelli, P. Salati, and P. D. Serpico, "AMS-02 antiprotons, at last! Secondary astrophysical component and immediate implications for dark matter," J. Cosmol. Astropart. Phys. 2015 (09), 023 (2015).
- M. Aguilar et al. (AMS Collab.), "Antiproton flux, antiproton-to-proton flux ratio, and properties of elementary particle fluxes in primary cosmic rays measured with the alpha magnetic spectrometer on the International Space Station," Phys. Rev. Lett. **117** (9) (2016).
- M. di Mauro, F. Donato, A. Goudelis, and P. D. Serpico, "New evaluation of the antiproton production cross section for cosmic ray studies," Phys. Rev. D 90 (8) (2014).
- M. W. Winkler, "Cosmic ray antiprotons at high energies," J. Cosmol. Astropart. Phys. 2017 (02), 048 (2017).
- F. Donato, M. Korsmeier, and M. D. Mauro, "Prescriptions on antiproton cross section data for precise theoretical antiproton UX predictions," Phys. Rev. D 96 (4) (2017).
- B. Adams et al. (COMPASS Collab.), "Letter of intent: A new QCD facility at the M2 beam line of the CERN SPS (COMPASS++/AMBER)," arXiv:1808.00848v6.
- I. Savin, A. Efremov, D. Peshekhonov, A. Kovalenko, O. Teryaev, O. Shevchenko, A. Nagajcev, A. Guskov, V. Kukhtin, and N. Toplilin, "Spin physics experiments

at NICA-SPD with polarized proton and deuteron beams," EPJ Web of Conf. **85**, 02039 (2015).

- V. Kekelidze, A. Kovalenko, R. Lednicky, V. Matveev, I. Meshkov, A. Sorin, and G. Trubnikov, "Status of the NICA project at JINR," EPJ Web of Conf. 138, 01027 (2017).
- 11. A. Guskov and R. El-Kholy, "On the possibility to study antiproton production at the SPD detector at NICA collider for dark matter search in astrophysical experiments," Phys. Part. Nucl. Lett. **16**, 216–223 (2019).
- A. Tkachenko, SpdRoot: The Framework for Simulation, Reconstruction and Analysis. http://spd.jinr.ru/wpcontent/uploads/2019/06/16Spd050619Tkachenko.pdf.
- M. Al-Turany, D. Bertini, R. Karabowicz, D. Kresan, P. Malzacher, T. Stockmanns, and F. Uhlig, "The Fair-Root framework," J. Phys.: Conf. Ser. **396**, 022001 (2012).
- T. Sjöstrand, S. Ask, J. R. Christiansen, R. Corke, N. Desai, P. Ilten, S. Mrenna, S. Prestel, C. O. Rasmussen, and P. Z. Skands, "An introduction to PYTHIA 8.2," Comput. Phys. Commun. 191, 159–177 (2015).
- 15. J. Adam et al. (ALICE Collab.), "Determination of the event collision time with the ALICE detector at the LHC," Eur. Phys. J. Plus **132** (2) (2017).
- R. Kappl and M. W. Winkler, "The cosmic ray antiproton background for AMS-02," J. Cosmol. Astropart. Phys. 2014 (09), 051 (2014).
- T. Sjöstrand, S. Mrenna, and P. Skands, "A brief introduction to PYTHIA 8.1," Comput. Phys. Commun. 178, 852–867 (2008).